



Topic E.1

STRUCTURE OF THE ATOM

*Every Formula, Trick & Trap on a Page
& IB-Style Quick Reference*

Topics Covered

The nuclear model
Geiger–Marsden–Rutherford
Atomic spectra
Nuclear radius & density (HL)
Bohr model of hydrogen (HL)

Exam Relevance

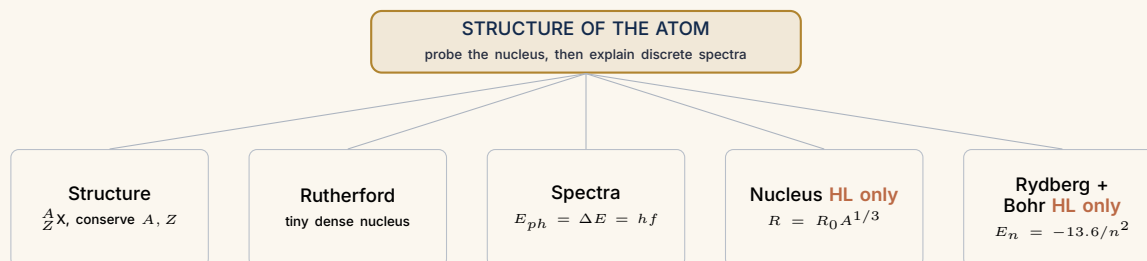
P1 & P2 (SL + HL)
Empty space + dense nucleus
Discrete photon transitions
 $R = R_0 A^{1/3}$
 $E_n = -13.6/n^2 \text{ eV}$

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The Big Picture — how E.1 fits together



1 Structure of Matter & the Nuclear Model

E.1 • SL + HL

Nuclear notation: A_ZX : A = nucleon number, Z = proton number, $N = A - Z$

Particle summary

Proton	charge $+e$	mass $\approx 1\text{ u}$	in nucleus
Neutron	charge 0	mass $\approx 1\text{ u}$	in nucleus
Electron	charge $-e$	mass $\approx m_p/1836$	orbits nucleus

Isotopes: same Z , different A (same element, different neutron count). $1\text{ u} = 1.66 \times 10^{-27}\text{ kg} = \frac{1}{12}$ mass of a ${}^{12}_6\text{C}$ atom.

TRICK In any nuclear reaction, conserve A (top) and Z (bottom) **separately**. A γ -photon carries no A or Z .

NOTE A classical orbiting (accelerating) electron should radiate and spiral in — atoms are stable, so electron energies must be **quantised** (§3, §7).

FROM THE PHOTON QUESTION BANK

Unified mass unit is near-guaranteed recall ($\frac{1}{12}$ of carbon-12 — not a proton/neutron average, not oxygen-16). "Given p/n/e, write A_ZX ": $A = p+n$, $Z = p$; electrons change *charge*, not notation. Equal-mass MCQ: smallest molar mass $\Rightarrow$ most atoms ($N \propto 1/M$).

2 Geiger–Marsden–Rutherford Evidence

E.1 • SL + HL

Setup & observations

Setup: α source $\rightarrow$ collimator $\rightarrow$ thin **gold foil** $\rightarrow$ ZnS screen, in vacuum.

Observation
Most α pass straight through
Few deflected $> 90^\circ$
 ~ 1 in 8000 back-scatter

Conclusion
atom is mostly **empty space**
positive charge & mass in a tiny nucleus
nucleus is very small, dense, positive

Number vs angle follows an **inverse-square** (Coulomb) law $\Rightarrow$ disproved the plum-pudding model.

TRAP Geiger–Marsden gives **no** evidence for discrete energy levels — only empty space + dense positive nucleus. Energy-level evidence comes from *spectra*.

NOTE Scale: atom $\sim 10^{-10}\text{ m}$, nucleus $\sim 10^{-15}\text{ m}$, ratio $\sim 10^5$ ("nucleus a football $\Rightarrow$ atom 30 km across").

FROM THE PHOTON QUESTION BANK

Most-repeated trap in the topic. Also: large-angle back-scatter is what *disproved plum-pudding*; answer observation $\rightarrow$ conclusion in pairs.

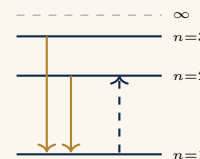
3 Atomic Spectra & Photon Transitions

E.1 • SL + HL

Photon energy: $E = hf = \frac{hc}{\lambda}$, $h = 6.63 \times 10^{-34}\text{ J s}$

Transition: $E_{\text{photon}} = E_{\text{upper}} - E_{\text{lower}}$ (emit ↓, absorb ↑)

Emission vs absorption & energy levels



Emission: hot gas $\Rightarrow$ bright discrete lines on dark.

Absorption: white light through cool gas $\Rightarrow$ dark lines at the **same** λ .

Discrete because electrons occupy only fixed levels $\Rightarrow$ only specific $\Delta E \Rightarrow$ specific λ .

Fingerprint: unique levels identify elements in stars & galaxies.

TRICK n levels $\Rightarrow \frac{n(n-1)}{2}$ lines. **Longest** λ = **smallest** gap. Two-step cascade: $\frac{1}{\lambda} = \frac{1}{\lambda_1} + \frac{1}{\lambda_2}$ (reciprocals add).

TRAP Electrons absorb photons of **exactly** the gap energy — any other energy passes straight through. *Excited* (higher level) $\neq$ *ionised* ($E \geq 0$).

FROM THE PHOTON QUESTION BANK

Energy-level diagrams appear almost every session — mind the negative signs in ΔE ; know count = $n(n-1)/2$, longest λ = smallest gap, and J $\leftrightarrow$ eV conversion.

4 Nuclear Radius & Density

E.1 • HL only

Nuclear radius: $R = R_0 A^{1/3}$, $R_0 = 1.2 \times 10^{-15} \text{ m} = 1.2 \text{ fm}$

Nuclear density: $\rho = \frac{m_u}{\frac{4}{3}\pi R_0^3} \approx 2.3 \times 10^{17} \text{ kg m}^{-3}$ (constant, all nuclei)

Ratio method

$\frac{R_Y}{R_X} = \left(\frac{A_Y}{A_X}\right)^{1/3}$. e.g. $A : 2 \rightarrow 16$ gives $8^{1/3} = 2$, so $R_Y = 2R_X$; density unchanged. Find A from R : $A = (R/R_0)^3$.

TRICK Density constant: $V \propto A$ and $m \propto A$, so the A cancels ($\rho = m_u / \frac{4}{3}\pi R_0^3$). Nucleons pack like identical marbles.

TRAP Never say density rises with A . It is $\sim$ constant $\approx 10^{14} \times$ ordinary matter.

FROM THE PHOTON QUESTION BANK

$R = R_0 A^{1/3}$ is the **most-examined HL formula** here — scale with the cube root or cube to find A ; “density constant” is a near-permanent distractor.

5 Closest Approach & Deviations from Rutherford

E.1 • HL only

Closest approach: $E_k = \frac{k q_\alpha q_{\text{nuc}}}{d} = \frac{k(2e)(Ze)}{d} \Rightarrow d = \frac{k(2e)(Ze)}{E_k}$

Deviations at high energy

Low energy: pure Coulomb $\Rightarrow$ Rutherford formula fits. **High energy:** α enters the range of the **strong nuclear force** $\Rightarrow$ fewer large-angle deflections than predicted.

Deviations $\Rightarrow$ (I) evidence for the strong force $\checkmark$ (II) estimate of nuclear radius $\checkmark$ (III) occur at *low* energy $\times$ (only HIGH).

TRICK At closest approach **all** KE $\rightarrow$ electric PE; rearrange $E_k = kq_1 q_2 / d$ for d .

TRAP d is from the nucleus **centre**; if $d \gg R$, Rutherford still holds.

FROM THE PHOTON QUESTION BANK

Two marks routinely lost: α charge is $2e$, and convert MeV $\rightarrow$ J before substituting. "Deviations at low energy" is the standard wrong option.

6 Hydrogen Spectrum & the Rydberg Formula

E.1 · HL only

Rydberg formula: $\frac{1}{\lambda} = R_H \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$, $R_H = 1.097 \times 10^7 \text{ m}^{-1}$, $n > n_1$

Spectral series (by lower level n_1)

Lyman $n_1=1$ (UV) **Balmer** $n_1=2$ (visible) **Paschen** $n_1=3$ (IR) Brackett/Pfund $n_1=4/5$ (IR)

NOTE Rydberg and Bohr describe the *same* lines: substitute $E_n = -13.6/n^2$ into $1/\lambda = (E_{\text{up}} - E_{\text{low}})/hc$ to recover Rydberg.

FROM THE PHOTON QUESTION BANK

Know the series by region; "which is visible?" $\rightarrow$ Balmer. Watch the common-denominator algebra and significant figures.

7 The Bohr Model of Hydrogen

E.1 · HL only

Energy levels: $E_n = \frac{-13.6}{n^2} \text{ eV}$

Angular momentum: $mvr = \frac{n\hbar}{2\pi} = n\hbar$

Radius / speed: $r_n = n^2 a_0$, $a_0 = 5.29 \times 10^{-11} \text{ m}$; $v_n \propto 1/n$

Hydrogen levels (memorise) & ratios

n	E_n	state
1	-13.6 eV	ground
2	-3.40 eV	1st excited
3	-1.51 eV	2nd excited
∞	0	ionised

$$L_n = n\hbar \Rightarrow L_2/L_1 = 2$$

$$r_n \propto n^2 \Rightarrow r_2/r_1 = 4$$

$$v_n \propto 1/n \Rightarrow v_2/v_1 = \frac{1}{2}$$

$$\text{Ionisation from } n: |E_n| = 13.6/n^2 \text{ eV}$$

TRAP Bohr works **only** for hydrogen / one-electron ions (He^+ , Li^{2+}). Never use $-13.6/n^2$ for neutral helium or heavier.

FROM THE PHOTON QUESTION BANK

Angular momentum quantised $L = n\hbar/2\pi$ (ratios collapse to a ratio of n). A favourite MCQ removes one electron from an isotope and asks if $-13.6/n^2$ applies; and to identify which listed equations are genuinely Bohr-model.

Exam-readiness self-check

Concept	Level	Confidence
Structure, notation, ${}^A_Z\text{X}$, isotopes, u	SL+HL	☐ ☐ ☐ ☐ ☐
Rutherford evidence & conclusions	SL+HL	☐ ☐ ☐ ☐ ☐
Spectra, energy levels, $E = hf$	SL+HL	☐ ☐ ☐ ☐ ☐
Nuclear radius & constant density	HL	☐ ☐ ☐ ☐ ☐
Closest approach & deviations	HL	☐ ☐ ☐ ☐ ☐
Hydrogen spectrum & Rydberg	HL	☐ ☐ ☐ ☐ ☐
Bohr model & ratios	HL	☐ ☐ ☐ ☐ ☐